Inverse Rendering of Lambertian Surfaces Using Subspace Methods

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Abstract—We propose a vector space approach for inverse rendering of a Lambertian convex object with distant light sources. In this problem, the texture of the object and arbitrary lightings are both to be recovered from multiple images of the object and its 3D model. Our work is motivated by the observation that all possible images of a Lambertian object lie around a low-dimensional linear subspace spanned by the first few spherical harmonics. The inverse rendering can therefore be formulated as a matrix factorization, in which the basis of the subspace is encoded in a spherical harmonic matrix S associated with the object's geometry. A necessary and sufficient condition on S for unique factorization is derived with an introduction to a new notion of matrix rank called nonseparable full rank. A singular value decomposition-based algorithm for exact factorization in the noiseless case is introduced. In the presence of noise, two algorithms, namely, alternating and optimization based are proposed to deal with two different types of noise. A random sample consensus-based algorithm is introduced to reduce the size of the optimization problem, which is equal to the number of pixels in each image. Implementations of the proposed algorithms are done on a real data set.

Index Terms—Computational relighting, Lambertian surfaces, inverse rendering, reflectance function, spherical convolution, spherical harmonics, matrix factorization, singular value decomposition, convex optimization.

I. INTRODUCTION

OMPUTATIONAL relighting is the problem of rendering images under virtual and novel lighting conditions. While it has been widely used on synthetic scenes constructed from computer graphics models or on real scenes captured under controlled lighting conditions, efficient relighting on general real scenes remain a research challenge due to its complexity involving both the geometric and photometric aspects in the image formation process. Existing relighting methods (see [1], [2] and the references therein) often consist in estimation, or inverse rendering of the scene geometry and texture,

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the BRDF (*Bi-directional Reflectance Distribution Function*) that defines the reflectance characteristics of the scene, and the lighting parameters. Given only the image data, the estimation of all these factors is ill-posed, and can only be achieved by strict assumptions on the scene geometry and/or light sources, or by controlled experiments (see [3]). The introduction of depth cameras and recent development of depth sensor fusion algorithms such as *KinectFusion* [4] open up new opportunities for efficient and reliable acquisition of 3D scene geometry in real-time. In this work, we explore the incorporation of geometry information in an inverse rendering problem under general, arbitrary (or *uncontrolled*) distant illumination.

Our work focuses on inverse rendering of *convex* Lambertian surfaces with *known geometry* given several images under various *distant* lighting conditions. In this model, the problem can be formulated, with the aid of a signal processing framework [5], [6], as a matrix factorization, in which an image matrix has to be decomposed into a product of an *albedo* diagonal matrix with a known *spherical harmonic* (SH) matrix in the middle and a lighting matrix. To the best of our knowledge, our work is the first one to study the problem of recovering both *nonuniform* albedos and uncontrolled lightings at the same time based on images and scene geometry. This is, to some extent, similar to an *array signal processing* problem [7], where both the DOAs (direction of arrivals) and sources are to be estimated from multiple snapshots of the array.

The main contribution of this paper is to solve the special matrix factorization mentioned above in both noiseless and noisy cases using the power of subspace methods. In the noiseless case, when both the lighting model and measurement are accurate, we gives a necessary and sufficient condition for unique factorization with an introduction to a new notion of matrix full rank, called nonseparable full rank (NSFR). We also solve for an exact solution using singular value decomposition (SVD) when the factorization is unique. In the noisy case we deal with the approximation error and measurement noise separately. When the approximation error is present, we propose an SVD-based algorithm to find the least squares solution of the factorization. This algorithm can be slightly modified to become a convex optimization that incorporates the positivity constraint on albedos. Another algorithm based on RANdom SAmple Consensus (RANSAC) [8] is also proposed to reduce the large size of the optimization problem which is equal to the number of pixels in each image. On the other hand, when the measurement noise is present, we propose an alternating least squares algorithm to find a local optimizer for the factorization.

The remainder of the paper is structured as follows. Sec. II reviews some of the related work. Sec. III discusses basic background of the reflection equation and spherical harmonic expansions. Sec. IV formulates the inverse rendering as a special matrix factorization problem. Sec. V studies the uniqueness and exact solution of the factorization. Sec. VI provides various algorithms for the matrix factorization in the presence of noise based on least squares and convex optimization. Sec. VII discusses a computational issue on the large size of the optimization problem and a heuristic way to overcome it. Sec. VIII presents some experimental results on a real dataset using the proposed algorithms. Finally, Sec. IX makes some concluding remarks.

II. RELATED WORK

The difficulties of inverse rendering under uncontrolled lighting are due to the lack of a discretized framework that can efficiently describe the reflectance function which has been often interpreted as a spherical integral. Ramamoorthi and Hanrahan in their series of work [5], [9], [10] introduced a signal processing framework for inverse rendering, in which a reflectance function is treated simply as a spherical convolution of a lighting function with the reflection kernel (BRDF multiplied by the half-cosine) and scaled by the albedos. This allows us to relate a reflectance function of an object to the lighting and the reflection kernel in terms of their SH expansions (sphere counterpart of Fourier series). Transforming from space-domain to frequency-domain yields two great advantages: (1) integrals are mapped to products of coefficients; and (2) reflectance function can be well-approximated by a few low-frequency terms (because reflection kernels often vary slowly). Based on this framework, the authors developed various algorithms for inverse rendering, most of which are under the form of an alternating optimization in which no global solution is guaranteed. Furthermore, they only considered homogeneous objects with no texture, and thus ignoring the local scalings by the albedos of the reflection. Our method deals with nonuniform albedos and can give exact solutions in the noiseless case.

Basri and Jacobs in an independent work [6] discovered a similar result for the special case of Lambertian (or diffuse) surfaces whose reflection kernels are simply the half-cosine function which contains mostly low-frequency components. As a result, it was shown in [6] that any reflectance function of a Lambertian *convex* homogeneous object with distant light sources can be well-approximated by its first 9 terms of the SH series. In other words, all the images of such an object live close to a 9-dimensional linear subspace spanned by the first 9 SHs associated with the geometry of the object. Basri et al., in the subsequent work [11] on photometric stereo (PS), matricized this key observation to formulate the inverse rendering as a matrix factorization which is similar to ours. In the PS problem, however, the albedos are often assumed to be uniform, and the scene geometry is unknown and to be reconstructed. In contrast, our work assumes known geometry and the nonuniform albedos are to be recovered. The authors of [11] proposed a matrix factorization based on the SVD followed by a low-rank approximation to estimate

the shape of the object. This method is very similar to Tomasi-Kanade factorization [12], that has been widely used in structure from motion, a classical problem in computer vision. However, this type of factorization is only within a linear ambiguity that cannot be fully resolved without any further prior knowledge. In particular, because the set of all possible images of an object, called its illumination cone, is invariant to any 3×3 linear transformations of the object's geometry [13], the 9-D subspace of SHs, and hence the factorization is approximately the same for different objects related by a linear transformation. If using only the first 4 SHs in the approximation, the linear ambiguity of the factorization can be reduced to a Lorentz transformation [14] with 7 degrees of freedom, by taking into account the unit norm of surface normals. Del Bue et al. [15] presented a bilinear factorization method to solve the PS problem, where one of the factors is constrained to lie on a specific manifold. The formulation decouples the core bilinear aspect from the manifold specificity and can be solved using Augmented Lagrange Multipliers. However, the manifold projector based on unit norm of surface normals only permits the use of first

More recently, Nießner, Zollhöfer, Wu et al. [16]–[18] introduced several methods for real-time inverse rendering from color plus depth data, in which the geometry is obtained from depth maps and some geometry refinement procedure. Unlike our work, these papers focus more on the practical side of the inverse rendering including geometry acquisition. However, similar to PS literature, these methods assume uniform albedos while estimating the lightings. In contrast, our paper puts more emphasis on the theoretical side of the matrix factorization problem where the lightings and albedos are estimated simultaneously using subspace methods.

For glossy objects when the Lambertian assumption is no longer hold, BRDF recovery has also been widely investigated in the computer graphics community [19]–[22]. In a recent work, Goldman *et al.* [23] estimated both geometry and BRDF from a small number of photographs. However, all of these methods for BRDF acquisition heavily rely on a controlled lighting setup in which the light can be easily calibrated. In this paper, we assume Lambertian objects with constant BRDF. This strict assumption allows us to rigorously tackle the inverse rendering under arbitrary general lighting conditions.

III. PRELIMINARIES

This section summarizes the signal-processing framework [1], [6] that relates a reflectance function to the lighting and reflection kernel as a spherical convolution in spacedomain. The relation can be transformed into frequency-domain via spherical harmonic expansions.

A. Reflection as Convolution

Throughout this paper, we restrict our interest to the illumination of Lambertian convex objects with distant light sources. The convexity assumption makes sure that there are no shadowing and inter-reflection which are very hard to analyze. Likewise, the Lambertian assumption implies that

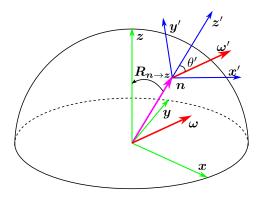


Fig. 1. Global and local coordinate systems. Green vectors define the global Cartesian coordinate system xyz. Blue vectors define a local coordinate system x'y'z' associated with the normal n such that z' = n. The light vector in red is labeled ω w.r.t. global coordinates and ω' w.r.t. local coordinates; θ' is the polar angle of ω' . The lighting function $L(\omega)$ is given in global, whereas the reflection kernel $K(\omega')$ is given in local coordinates. The two coordinates are related by the rotation that transfers the normal n (or z') to the polar vector z: $\omega' = R_{n \to z} \omega$.

the reflection does not depend on the viewing angle, and so significantly simplifies our lighting model. We want to remark that, although assumed to be distant, the light source can be arbitrarily complex but not necessarily a point source.

In the following analysis, we refer to Fig. 1 that illustrates the *global* and *local* coordinate systems associated with a location on the surface of the object. Let p be a point on the surface with normal vector n, and $L(\omega)$ be the lighting function that defines the incident radiance at angle ω w.r.t. some global coordinate system xyz. The reflected radiance (or *reflectance function*) Y(p,n) at location p and normal n, is then given by integrating the reflections over all possible light directions. In particular,

$$Y(\mathbf{p}, \mathbf{n}) = \rho(\mathbf{p}) \int_{\Omega} L(\boldsymbol{\omega}) \cos(\theta') d\boldsymbol{\omega}', \tag{1}$$

where $\rho(p)$ is the albedo at location p; ω' is the light direction w.r.t. the local coordinate system x'y'z' obtained by rotating xyz so that z'=n; θ' is the corresponding polar angle of ω' ; and the integrating region Ω is the upper-hemisphere of the unit sphere. The integral in (1) can be extended to the whole unit sphere S as

$$Y(\mathbf{p}, \mathbf{n}) = \rho(\mathbf{p}) \int_{S} L(\boldsymbol{\omega}) K(\boldsymbol{\omega}') d\boldsymbol{\omega}', \qquad (2)$$

where $K(\omega') \stackrel{\triangle}{=} \max(\cos\theta', 0)$ is a *half-cosine* function of the local polar angle θ' . We will later refer to $K(\omega')$ as the *Lambertian kernel*. Note that, although written as $K(\omega')$ for general purpose, this kernel is actually a 1-D function that depends only on the polar angle of the incident direction in local coordinates.

By definition, the global and local coordinates are related by a 3D rotation: $\omega' = R_{n\to z}\omega$, where $R_{n\to z}$ is the rotation matrix that brings n=z' back to z. Substituting this relation into (2), we arrive at

$$Y(\mathbf{p}, \mathbf{n}) = \rho(\mathbf{p}) \int_{S} L(\boldsymbol{\omega}) K(\mathbf{R}_{\mathbf{n} \to \mathbf{z}} \boldsymbol{\omega}) d\boldsymbol{\omega}, \tag{3}$$

We can easily recognize the integral in (3) as a *spherical* convolution, in which the translation operator in a classical

circular convolution is replaced with a rotation operator. We adopt the notation \circledast_s for spherical convolution and write

$$(L \circledast_s K)(n) \stackrel{\Delta}{=} \int_S L(\omega) K(\mathbf{R}_{n \to z} \omega) d\omega. \tag{4}$$

Similar to circular convolutions, the above spherical convolution defines a *linear rotation-invariant* (LRI) system with impulse response $K(\omega)$.

To summarize the above analysis, the reflectance function is now a spherical convolution of the lighting and Lambertian kernel scaled by albedos, i.e.,

$$Y(\mathbf{p}, \mathbf{n}) = \rho(\mathbf{p})(L \circledast_{s} K)(\mathbf{n}). \tag{5}$$

B. Spherical Harmonics as Basis

Now we want to look at the reflection equation (5) in frequency-domain because the energy of the Lambertian kernel $K(\omega')$ is compacted in low frequencies. As a spherical counterpart of Fourier basis on the circle, the *spherical harmonics* $\{S_{m,n}\}_{m\geq 0, |n|\leq m}$ form an orthonormal basis for functions on the unit sphere, whose formulas are given by

$$S_{m,n}(\boldsymbol{\omega}) = N_{m,n} \cdot P_{m,n}(\cos \theta) e^{jn\varphi}, \tag{6}$$

where $j = \sqrt{-1}$; θ and φ are respectively the polar and azimuthal angles of ω : $\theta = \arccos(z)$, $\varphi = \arctan(y/x)$; $P_{m,n}(\cdot)$ is the associated Legendre function; and $N_{m,n}$ is the normalized term.¹

Similarly to Fourier series of functions on a circle, every function $F(\omega)$ on the unit sphere has a spherical harmonic (SH) expansion

$$F(\boldsymbol{\omega}) = \sum_{m=0}^{\infty} \sum_{n=-m}^{m} f_{m,n} \cdot S_{m,n}(\boldsymbol{\omega}),$$

where the SH coefficients $f_{m,n}$ are given by

$$f_{m,n} = \int_{S} F(\boldsymbol{\omega}) S_{m,n}^{*}(\boldsymbol{\omega}) d\boldsymbol{\omega}.$$

The feature that makes spherical harmonics similar to Fourier basis on a circle is that they are eigensignals of LRI systems.² This key property maps a spherical convolution in space-domain to a multiplication in frequency-domain. Particularly, let the SH coefficients of L, K and $L \circledast_s K$ be $\{\ell_{m,n}\}$, $\{k_{m,n}\}$, and $\{x_{m,n}\}$, respectively, then it was shown in [5] and [6] that

$$x_{m,n} = \sqrt{\frac{4\pi}{2m+1}} \, k_{m,0} \, \ell_{m,n}, \quad m \ge 0, |n| \le m \tag{7}$$

Moreover, it was also shown in [6] that the convolution $L \otimes_s K$ can be well-approximated using first few terms in its SH expansion, namely

$$(L \circledast_s K)(n) \approx \sum_{m'=0}^{M'-1} \sum_{n'=-m'}^{m'} x_{m',n'} \cdot S_{m',n'}(n),$$

¹See [6] for specific formulas of $P_{m,n}(\cdot)$ and $N_{m,n}$.

²See a special case of the Funk-Hecke theorem [6, Th. 1]. Note that since $K(\theta')$ is a 1-D function of the polar angle, its SH coefficients $k_{m,n}$ are nonzero only for n=0.

for small M' (for instance, using M' = 3 preserves roughly 97.96% the energy of $L \circledast_s K$.) As a result, the reflectance function can be approximated by

$$Y(\boldsymbol{p}, \boldsymbol{n}) \approx \rho(\boldsymbol{p}) \sum_{m=1}^{M} x_m \cdot S_m(\boldsymbol{n}),$$
 (8)

where $M = M'^2$, and the double index (m', n') was converted to the single index $m = m'^2 + m' + n' + 1$. In words, this approximation says that, all the reflectance functions of a Lambertian textureless object live around a low-dimensional linear subspace spanned by the first few SHs evaluated at the normal vectors of the object.

It is important to note that the SHs presented above are complex-valued (for analytical simplicity). In experiments, however, we can use the real-valued SHs which are real and imaginary parts of the above functions. From now on, we assume that the SHs, $\{S_m(\omega)\}_{m\geq 1}$, are real-valued, where the specific forms for the first 9 of them at $\omega=(x,y,z)$, for $x^2+y^2+z^2=1$, are given below [6, eq. (7)]

$$S_{1} = \frac{1}{\sqrt{4\pi}}, \quad S_{2} = \sqrt{\frac{3}{4\pi}}x, \quad S_{3} = \sqrt{\frac{3}{4\pi}}y, \quad S_{4} = \sqrt{\frac{3}{4\pi}}z,$$

$$S_{5} = \sqrt{\frac{15}{4\pi}}xy, \quad S_{6} = \sqrt{\frac{15}{4\pi}}yz, \quad S_{7} = \sqrt{\frac{5}{16\pi}}(3z^{2} - 1),$$

$$S_{8} = \sqrt{\frac{15}{4\pi}}zx, \quad S_{9} = \sqrt{\frac{15}{16\pi}}(x^{2} - y^{2}). \tag{9}$$

IV. PROBLEM STATEMENT

A. From Reflection to Images

Suppose we are given J images taken at the same viewpoint of the same object under different lighting conditions. Let N be the number of pixels in each image, and p_i be the location on the surface corresponding to pixel i on each image, for i = 1, ..., N. From (8), the intensity at pixel i of image j can be approximated by

$$Y_j(\boldsymbol{p}_i, \boldsymbol{n}_i) \approx \rho(\boldsymbol{p}_i) \sum_{m=1}^{M} x_{m,j} \cdot S_m(\boldsymbol{n}_i), \tag{10}$$

for $1 \le i \le N$, $1 \le j \le J$. We can put all equations in (10) into a single matrix form by defining matrices $Y \in \mathbb{R}^{N \times J}$; $\Phi \in \mathbb{D}^N$ (the set of all size-N diagonal matrices); $X \in \mathbb{R}^{M \times J}$; and $S \in \mathbb{R}^{N \times M}$ as

$$Y_{ij} = Y_j(\mathbf{p}_i), \quad \Phi_{ii} = \rho(\mathbf{p}_i), \quad X_{mj} = x_{m,j},$$

and

$$S_{im} = S_m(\mathbf{n}_i), \tag{11}$$

for $1 \le i \le N$, $1 \le j \le J$, $1 \le m \le M$, and M_{ij} denotes the entry of matrix M at row i and column j. With these notations, (10) becomes

$$Y \approx \Phi S X$$
, (12)

where Y, Φ , S and X will be respectively referred to as *image*, albedo, spherical harmonic and lighting matrices.

B. The Matrix Factorization

For analytical purpose, we first assume the noiseless case when the approximation in (12) is replaced with exact equation

$$Y = \Phi S X, \tag{13}$$

In light of (13), the inverse rendering problem becomes a matrix factorization in which the image matrix $Y \in \mathbb{R}^{N \times J}$ and SH matrix $S \in \mathbb{R}^{N \times M}$ are known (can be computed from given images and 3D model of the object); the albedo matrix $\Phi \in \mathbb{D}^N$ and lighting matrix $X \in \mathbb{R}^{M \times J}$ are to be recovered. Once the albedo and lighting matrices are reconstructed, say $\hat{\Phi}$ and \hat{X} , the forward rendering becomes an obvious matrix multiplication

$$y_{\text{new}} = \hat{\Phi} S x_{\text{new}},$$

where y_{new} corresponds to a novel image under a virtual lighting generated by a vector x_{new} . The rest of the paper is dedicated to the problem of solving (Φ, X) from (13) given (Y, S), which will be referred to as the matrix factorization.

We first note this matrix factorization must always have a solution $(\Phi_0, X_0) \in \mathbb{D}^N \times \mathbb{R}^{M \times J}$, where \mathbb{D}^N denotes the set of all $N \times N$ diagonal matrices, as the image matrix Y is collected according to the above lighting model. The recovery of (Φ, X) is possible only if (Φ_0, X_0) is the unique solution of the matrix factorization, within a scaling. Therefore, we need to assume further that: (1) $\Phi_0 \in \mathbb{D}^N_*$, or all diagonal entries of Φ_0 are nonzero, and (2) $X_0 \in \mathbb{R}^{M \times J}_{\text{full}}$, or X_0 has full row rank. The need for these assumptions will be justified later in the proof of uniqueness. Here, we just want to make a few remarks on them. The first assumption that all albedos are nonzero, does not restrict ourselves because the zero-intensity pixels corresponding to zero albedos can be masked out of the equations. Likewise, the second assumption that X_0 has full row rank is reasonable and often made in classical algorithms for array signal processing such as MUSIC (MUltiple SIgnal Classification) [24] and ESPRIT (Estimation of Signal Parameters via Rotational Invariance Techniques) [25]. It essentially means that the images are to be taken under diversified lighting conditions. This assumption also implicitly requires that the number of images J must be greater than the dimension M of the approximation subspace.

V. Noiseless Case: Uniqueness and Exact Solution

We start with the noiseless case when (13) holds, ignoring both the approximation error and measurement noise. The first question one should ask when dealing with an inverse problem is when the recovery is unique. This section provides a necessary and sufficient condition on the SH matrix S such that the matrix factorization (13) is unique (up to a scaling factor), and an SVD-based method for recovery.

A. Uniqueness

Before stating the condition for uniqueness, we introduce some new notions of matrix full rank (FR) which will be useful later. The results of this subsection might be also of independent interests. Consider a tall matrix $S \in \mathbb{R}^{N \times M}$

with N > M. Let \mathcal{N} denote the set $\{1, 2, ..., N\}$. For any subset \mathcal{J} of \mathcal{N} , let \mathcal{J}^c be the complement of \mathcal{J} in \mathcal{N} , and let $S_{\mathcal{J}}$ be the submatrix constituted from the rows of S indexed by \mathcal{J} . We say \mathcal{J} is a nontrivial subset of \mathcal{N} if $\mathcal{J} \neq \mathcal{N}$ and $\mathcal{J} \neq \emptyset$.

Definition 1 ([26], [27]): The Kruskal rank of S, denoted by krank(S) is the maximal number k such that any k rows of S are linearly independent. S is said to have Kruskal full rank (KFR) if krank(S) = M.

Definition 2: **S** is said to have strong full rank (SFR) if it has no zero rows and rank $(S_{N\setminus\{i\}}) = M$ for all $i \in N$.

Definition 3: S is said to have nonseparable full rank (NSFR) if it has no zero rows, and rank(S) = M, and there does not exist a nontrivial subset \mathcal{J} of \mathcal{N} such that rank($S_{\mathcal{J}}$) + rank($S_{\mathcal{J}^c}$) = M.

In words, a matrix has NSFR if it has full column rank and its rows cannot be separated into two groups whose ranks add to the rank of the matrix. That justifies the term "nonseparable full rank." The following result gives another description of NSFR.

Proposition 1: **S** has NSFR if and only if $\operatorname{rank}(S_{\mathcal{J}}) + \operatorname{rank}(S_{\circ,\mathcal{J}}) \geq M+1$, for every nontrivial subset \mathcal{J} of \mathcal{N} .

Proof: Suppose that S has NSFR, then for every nontrivial subset $\mathcal J$ of $\mathcal N$ we have

$$rank(\mathbf{S}_{\mathcal{J}}) + rank(\mathbf{S}_{\mathcal{J}^c}) \neq M.$$

On the other hand

$$\operatorname{rank}(S_{\mathcal{T}}) + \operatorname{rank}(S_{\mathcal{T}^c}) \ge \operatorname{rank}(S) = M.$$

Hence

$$rank(\mathbf{S}_{\mathcal{J}}) + rank(\mathbf{S}_{\mathcal{J}^c}) \ge M + 1, \quad \forall \ \emptyset \ne \mathcal{J} \subsetneq \mathcal{N} \quad (14)$$

completing the "only if" part. For the "if" part, suppose (14) holds, then it is obvious that $\operatorname{rank}(S_{\mathcal{J}}) + \operatorname{rank}(S_{\mathcal{J}^c}) \neq M$, for every nontrivial subset \mathcal{J} of \mathcal{N} . We only need to show that S has no zero rows and $\operatorname{rank}(S) = M$. Indeed, for every $i \in \mathcal{N}$ we have

$$\operatorname{rank}(\mathbf{S}_{\{i\}}) + \operatorname{rank}(\mathbf{S}_{\mathcal{N}\setminus\{i\}}) \ge M + 1. \tag{15}$$

Since $\operatorname{rank}(S_{\{i\}}) \leq 1$ and $\operatorname{rank}(S_{\mathcal{N}\setminus\{i\}}) \leq \operatorname{rank}(S) \leq M$, (15) implies that $\operatorname{rank}(S_{\{i\}}) = 1$, and $\operatorname{rank}(S_{\mathcal{N}\setminus\{i\}}) = M$. From $\operatorname{rank}(S_{\{i\}}) = 1$, $\forall i \in \mathcal{N}$, we can deduce that S has no zero rows. Meanwhile, $\operatorname{rank}(S_{\mathcal{N}\setminus\{i\}}) = M$ implies that $\operatorname{rank}(S) = M$, completing the "if" part.

The definition of NSFR is not very intuitive. The following fact gives us a sense of how strong NSFR is, in comparison to KFR and SFR, whose definitions are more intuitive.

Proposition 2: The following implications hold: $KFR \Rightarrow NSFR \Rightarrow SFR \Rightarrow FR$. The reverse implications are generally not true. However, for N = M + 1, KFR, NSFR, and SFR are equivalent.

Fig. 2 illustrates the four concepts of full rank in the relations described in Proposition 2. We would like to emphasize

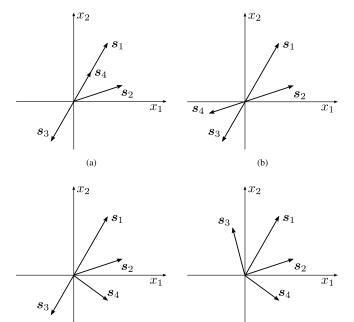


Fig. 2. Illustration of different notions of full rank for M=2, N=4 and $S=[s_1,s_2,s_3,s_4]^T$. (a) Full rank. (b) Strong full rank. (c) Nonseparable full rank. (d) Kruskal full rank.

(c)

(d)

that the number of rows of S in our matrix factorization problem is typically large, resulting in costly computations. The following result will be useful later when we want to reduce the size of the problem.

Proposition 3: **S** has NSFR if there exists a subset \mathcal{J} of \mathcal{N} such that $\mathbf{S}_{\mathcal{J}}$ has NSFR.

Proof: See Appendix B
$$\Box$$

With the NSFR carefully defined, we are ready to state the necessary and sufficient condition for the uniqueness of the matrix factorization up to a scaling factor. This is certainly the best we can hope for because if (Φ_0, X_0) is a solution then so is $(\alpha \Phi_0, \frac{1}{\alpha} X_0)$, for any scalar $\alpha \neq 0$. In the remainder of the paper, we frequently use the \odot symbol to denote point-wise operators between matrices. Namely,

$$m{A} \odot m{B} \stackrel{\Delta}{=} \{a_{ij} \cdot b_{ij}\}_{i,j},$$

 $m{A} \odot / m{B} \stackrel{\Delta}{=} \{a_{ij}/b_{ij}\}_{i,j},$
 $m{A} \odot^{\alpha} \stackrel{\Delta}{=} \{a_{ij}^{\alpha}\}_{i,j}.$

Theorem 1: Given a pair of matrices (Y, S), the matrix factorization (13) has no solutions $(\Phi, X) \in \mathbb{D}^N \times \mathbb{R}^{M \times J}$ other than $(\alpha \Phi_0, \frac{1}{\alpha} X_0)$, for some scalar $\alpha \neq 0$, if and only if S has NSFR.

Proof: For brevity, we only show here the proof for the sufficiency part; the proof for the necessity part is in Appendix E.

Suppose *S* has NSFR and there exist $(\tilde{\Phi}, \tilde{X}) \in \mathbb{D}^N \times \mathbb{R}^{M \times J}$ such that

$$Y = \Phi_0 S X_0 = \tilde{\Phi} S \tilde{X}. \tag{16}$$

We will show that $(\tilde{\Phi}, \tilde{X}) = (\alpha \Phi_0, \frac{1}{\alpha} X_0)$ for some $\alpha \neq 0$. Assume conversely that $\tilde{\Phi}$ is not a scaled version of Φ_0 . We first note that, by definition, S has full column rank,

 $^{^3}$ Kruskal rank is typically defined for fat matrices, requiring linear independence of any k columns. Here, we define it for tall matrices to conform to the context of the matrix factorization problem stated in Section IV.

i.e., $\operatorname{rank}(S) = M$. Let $\Phi_0 = \operatorname{diag}(\varphi)$, and $\tilde{\Phi} = \operatorname{diag}(\tilde{\varphi})$, where $\varphi \in \mathbb{R}^N_*$ and $\tilde{\varphi} \in \mathbb{R}^N_*$. Since S has no zero rows and X_0 has full row rank, each row of the product matrix SX_0 must be different from zero vector. Furthermore, since $\varphi_i \neq 0, \forall i \in \mathcal{N}$, each row of Y must also be different from zero vector. From this and (16), we can deduce that $\tilde{\varphi}_i \neq 0, \forall i \in \mathcal{N}$.

It is then valid to put $\psi = \varphi \odot /\tilde{\varphi}$, and $\Psi = \operatorname{diag}(\psi)$. From (16) we get

$$\Psi S X_0 = S \tilde{X}. \tag{17}$$

Since X_0 has full row rank, (17) is equivalent to

$$\Psi S = S\tilde{X}X_0^{\dagger},\tag{18}$$

where $X_0^{\dagger} = X_0^T (X_0 X_0^T)^{-1}$ is the pseudo-inverse of X_0 . Consider the following matrix

$$\hat{S} = [S|\Psi S] = \begin{bmatrix} s_1^T & \psi_1 s_1^T \\ s_2^T & \psi_2 s_2^T \\ \vdots & \ddots & \vdots \\ s_N^T & \psi_N s_N^T \end{bmatrix}, \tag{19}$$

where $s_i^T \in \mathbb{R}^M$ be the *i*-th row vector of S for $i \in \mathcal{N}$. Before proceeding we need the following lemmas whose proofs can be found in Appendices C and D.

Lemma 1: $\operatorname{rank}(\hat{S}_{\mathcal{I}}) = \operatorname{rank}(S_{\mathcal{I}}), \text{ for all } \mathcal{I} \subset \mathcal{N}.$ Especially, $\operatorname{rank}(\hat{S}) = \operatorname{rank}(S) = M.$

Lemma 2: For any $\mathcal{I} \subset \mathcal{N}$, if \mathcal{B} is a subset of \mathcal{I} such that $\{s_i\}_{i \in \mathcal{B}}$ is a basis for $\mathcal{R}(S_{\mathcal{I}}^T)$ then $\{[s_i^T | \psi_i s_i^T]^T\}_{i \in \mathcal{B}}$ is a basis for $\mathcal{R}(\hat{S}_{\mathcal{I}}^T)$.

Let us now continue with the proof of Thm. 1. The hypothesis that $\tilde{\Phi}$ is not a scaled version of Φ_0 implies that elements of the set $\{\psi_i\}_{i=1}^N$ are not all equal. Thus, there must exist a nontrivial subset \mathcal{J} of \mathcal{N} such that $\psi_j = \theta, \forall j \in \mathcal{J}$ and $\psi_j \neq \theta, \forall j \in \mathcal{J}^c$, for some fixed $\theta \neq 0$.

Now that S has NSFR, from Proposition 1 we must have

$$rank(S_{\mathcal{T}}) + rank(S_{\mathcal{T}^c}) \ge M + 1. \tag{20}$$

Combining with Lemma 1 yields

$$\operatorname{rank}(\hat{\mathbf{S}}_{\mathcal{J}}) + \operatorname{rank}(\hat{\mathbf{S}}_{\mathcal{J}^c}) \ge M + 1. \tag{21}$$

Expressing this in terms of dimensions gives

$$\begin{split} M+1 &\leq \dim\left(\mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}}^{T})\right) + \dim\left(\mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}^{c}}^{T})\right) \\ &= \dim\left(\mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}}^{T}) + \mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}^{c}}^{T})\right) + \dim\left(\mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}}^{T}) \cap \mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}^{c}}^{T})\right) \\ &= \dim\left(\mathcal{R}(\hat{\boldsymbol{S}}^{T})\right) + \dim\left(\mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}}^{T}) \cap \mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}^{c}}^{T})\right) \\ &= M + \dim\left(\mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}}^{T}) \cap \mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}^{c}}^{T})\right). \end{split}$$

Therefore

$$\dim \left(\mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}}^T) \cap \mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}^c}^T) \right) \geq 1.$$

There thus exits a nonzero vector $\hat{\mathbf{s}}$ in $\mathcal{R}(\hat{\mathbf{S}}_{\mathcal{J}}^{I}) \cap \mathcal{R}(\hat{\mathbf{S}}_{\mathcal{J}^{c}}^{I})$. Let $\mathcal{K} \subset \mathcal{J}$ and $\mathcal{L} \subset \mathcal{J}^{c}$ such that $\{\mathbf{s}_{k}\}_{k \in \mathcal{K}}$ is a basis of $\mathcal{R}(\mathbf{S}_{\mathcal{J}}^{T})$ and $\{\mathbf{s}_{\ell}\}_{\ell \in \mathcal{L}}$ is a basis of $\mathcal{R}(\mathbf{S}_{\mathcal{J}^{c}}^{T})$. It follows from Lemma 2 that $\{[\mathbf{s}_{k}^{T}|\theta\mathbf{s}_{k}^{T}]^{T}\}_{k \in \mathcal{K}}$ is a basis of $\mathcal{R}(\hat{\mathbf{S}}_{\mathcal{J}}^{T})$

and $\{[s_{\ell}^T | \psi_{\ell} s_{\ell}^T]^T\}_{\ell \in \mathcal{L}}$ is a basis of $\mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}^c}^T)$. Then $\hat{\boldsymbol{s}}$ can be represented in two different ways as

$$\hat{\mathbf{s}} = \sum_{k \in \mathcal{K}} \alpha_k [\mathbf{s}_k^T | \theta \mathbf{s}_k^T]^T = \sum_{\ell \in \mathcal{L}} \beta_\ell [\mathbf{s}_\ell^T | \psi_\ell \mathbf{s}_\ell^T]^T. \tag{22}$$

Hence

$$\sum_{k \in \mathcal{K}} \alpha_k s_k = \sum_{\ell \in \mathcal{L}} \beta_\ell s_\ell = \sum_{\ell \in \mathcal{L}} \frac{\beta_\ell \psi_\ell}{\theta} s_\ell. \tag{23}$$

By the linear independence of $\{s_\ell\}_{\ell\in\mathcal{L}}$ we have

$$\beta_{\ell} = \beta_{\ell} \psi_{\ell} / \theta, \quad \forall \ell \in \mathcal{L}.$$
 (24)

Since $\hat{s} \neq \mathbf{0}$, it follows from (22) that there exists a $\ell^* \in \mathcal{L}$ such that $\beta_{\ell^*} \neq 0$. This together with (24) yields $\psi_{\ell^*} = \theta$, contradicting to the above hypothesis that $\psi_j \neq \theta$ for all $j \in \mathcal{J}^c$. Therefore $\{\psi_i\}_{i=1}^N$ must be all equal, i.e. $\tilde{\Phi} = \alpha \Phi$ for some $\alpha \neq 0$. From (16), because S has full column rank, it must be that

$$\tilde{X} = S^{\dagger} \tilde{\Phi}^{-1} Y = \frac{1}{a} S^{\dagger} \Phi_0^{-1} Y = \frac{1}{a} X_0,$$
 (25)

where $S^{\dagger} = (S^T S)^{-1} S^T$ is the pseudo-inverse of S. This completes the proof of the sufficiency part.

B. Exact Solution

This subsection is to solve the matrix factorization (13) given that it has unique solution, i.e. S has NSFR by Thm. 1. Note that since S has full column rank, once Φ is known, S can be uniquely recovered by

$$X = S^{\dagger}(\Phi^{-1}Y). \tag{26}$$

Therefore, we can focus on finding Φ which has a diagonal structure. We propose an SVD-based algorithm to exactly recover Φ . The algorithm is based on the following theorem, our second main result in the noiseless case.

Theorem 2: $\Phi = \operatorname{diag}(\varphi)$ is a solution to (13) if and only if $\mathbf{z} \stackrel{\Delta}{=} \varphi \odot^{-1}$ is a nontrivial solution to

$$\left((I - SS^{\dagger}) \odot (YY^T) \right) z = \mathbf{0}. \tag{27}$$

Proof: Let $z \stackrel{\Delta}{=} \varphi \odot^{-1}$, then it is clear that φ solves $Y = \operatorname{diag}(\varphi)SX$ if and only if z solves

$$\operatorname{diag}(z)Y = SX. \tag{28}$$

It is equivalent to that each column of the LHS is in the range space of S, i.e.

$$(\operatorname{diag}(z)Y)_{j} \in \mathcal{R}(S), \quad \forall j = 1, \dots, J.$$
 (29)

Note that a vector lies in some subspace if and only if its orthogonal projection onto the subspace is equal to itself. Thus, (29) can be written equivalently as

$$P_{\mathcal{R}(S)}(\operatorname{diag}(z)Y)_j = (\operatorname{diag}(z)Y)_j, \quad \forall j = 1, \dots, J.$$

Replacing the projection matrix $P_{\mathcal{R}(S)}$ by SS^{\dagger} , we arrive at

$$(I - SS^{\dagger})(\operatorname{diag}(z)Y)_{i} = \mathbf{0}, \quad \forall i = 1, \dots, J.$$
 (30)

Algorithm 1 SVD-Based Factorization

Inputs: Y, S

Outputs: Estimates $\hat{\Phi}$, \hat{X} of Φ , X

- 1. Compute $M = (I SS^{\dagger}) \odot (YY^T)$.
- 2. Compute the eigenvalue decomposition of M.
- 3. Let z_{\min} be the eigenvector associated with the smallest eigenvalue of M.
- 4. Return $\hat{\mathbf{\Phi}} = \operatorname{diag}(\boldsymbol{z}_{\min} \odot^{-1}); \ \hat{\boldsymbol{X}} = \boldsymbol{S}^{\dagger} \operatorname{diag}(\boldsymbol{z}_{\min}) \boldsymbol{Y}.$

Let $\mathbf{Q} \stackrel{\Delta}{=} \mathbf{I} - \mathbf{S}\mathbf{S}^{\dagger}$. For each j = 1, ..., J, let $\mathbf{D}_i \stackrel{\Delta}{=}$ $diag((Y)_i)$, then we can rewrite (30) as

$$QD_{i}z = 0, \quad \forall j = 1, \dots, J. \tag{31}$$

Define matrix A by

$$\boldsymbol{A} \stackrel{\triangle}{=} \begin{bmatrix} \boldsymbol{Q} \boldsymbol{D}_1 \\ \vdots \\ \boldsymbol{Q} \boldsymbol{D}_J \end{bmatrix}. \tag{32}$$

With this notation, we can compact all J equations in (31) into a single one as

$$Az = \mathbf{0},\tag{33}$$

which is in turn equivalent to

$$A^T A z = \mathbf{0}. (34)$$

We can complete the proof at this point by using the following lemma whose proof is given in Appendix F.

Lemma 3: For matrix A defined in (32), we have

$$\mathbf{A}^T \mathbf{A} = (\mathbf{I} - \mathbf{S} \mathbf{S}^{\dagger}) \odot (\mathbf{Y} \mathbf{Y}^T).$$

Thm. 2 naturally gives rise to Alg. 1, in which we solve (27) for z by picking the eigenvector associated with the smallest eigenvalue of the positive definite matrix $(I - SS^{\dagger}) \odot (YY^{T})$. Once Φ is found, X can be computed using (26).

The following corollary is very useful in checking whether a matrix S has NSFR. It can be deduced directly from Thm. 1 and Thm. 2 by setting $\Phi_0 = I_N$, $X_0 = I_M$, and Y = S.

Corollary 1: A matrix $S \in \mathbb{R}^{N \times M}$ has NSFR if and only if

$$\operatorname{rank}\left((\boldsymbol{I} - \boldsymbol{S}\boldsymbol{S}^{\dagger}) \odot (\boldsymbol{S}\boldsymbol{S}^{T})\right) = N - 1$$

rank $(I - SS^{\dagger}) \odot (SS^T) = N - 1$. We conclude this subsection by giving a few comments on this result. Without identifying the NSFR with the uniqueness of the corresponding matrix factorization, we can hardly see the connection between Definition 3 and Corollary 1. When fixing M and growing N, checking if an $N \times M$ matrix has NSFR using the brute-force approach (i.e., computing the rank of every submatrix) would be exponentially complex. However, Corollary 1 provides a much more efficient indirect way to do so, in which only the rank of an $N \times N$ matrix needs to be computed, resulting in a polynomial complexity.

VI. NOISY CASE: ALGORITHMS

It is important to note that the equation in (13) can never be the case due to the approximation and measurement errors. Taking into account these noises, (13) should be remodeled as

$$Y = \Phi(SX + W_1) + W_2, \tag{35}$$

where W_1 and W_2 are respectively the approximation and measurement noises. Here, we assume both of them are deterministic unknown quantities. It is natural to pose the following least squares problem

$$\min \left\| \begin{bmatrix} \mathbf{W}_1 \\ \mathbf{W}_2 \end{bmatrix} \right\|_{\mathbf{E}} \text{ s.t. } \mathbf{Y} = \mathbf{\Phi}(\mathbf{S}\mathbf{X} + \mathbf{W}_1) + \mathbf{W}_2, \quad (P0)$$

where $\|\cdot\|_F$ denotes the Frobenius norm. Solving (P0) in the presence of both W_1 and W_2 is generally hard. In the sequel, we only tackle the two special cases of this noisy problem when either W_1 or W_2 is negligible.

A. No Measurement Error

If there is no measurement error, i.e. $W_2 = 0$, (35) can be rewritten as

$$\Phi^{-1}Y = SX + W_1.$$

Therefore, (P0) becomes

$$\min_{\substack{\boldsymbol{\Phi} \in \mathbb{D}_{X}^{N}, \|\boldsymbol{\Phi}^{-1}\|_{F} = 1 \\ \boldsymbol{Y} \in \mathbb{D}^{M \times J}}} \|\boldsymbol{\Phi}^{-1}\boldsymbol{Y} - \boldsymbol{S}\boldsymbol{X}\|_{F}, \quad (P1)$$

where the constraint $\|\mathbf{\Phi}^{-1}\|_F = 1$ is added to resolve the scaling ambiguity. The solution to this problem turns out to be identical to the exact solution in the noiseless case.

Theorem 3: The solution to the least squares problem (P1) is given by

$$\Phi_{LS} = \operatorname{diag}(z_{\min} \odot^{-1}), \tag{36}$$

and

$$X_{LS} = S^{\dagger} \operatorname{diag}(z_{\min}) Y, \tag{37}$$

where z_{\min} is the eigenvector associated with the smallest eigenvalue of $(I - SS^{\dagger}) \odot (YY^{T})$.

Proof: Fixing $\Phi \in \mathbb{D}_*^N$ such that $\|\Phi^{-1}\|_F = 1$, the least squares solution for X is given by

$$X_{LS}(\mathbf{\Phi}) = \arg\min_{X \in \mathbb{R}^{M \times J}} \|\mathbf{\Phi}^{-1} Y - SX\|_F$$
$$= S^{\dagger} \mathbf{\Phi}^{-1} Y. \tag{38}$$

Substituting into (P1) and using the matrix notations in Sec. V we can solve for the least squares solution of Φ as

$$\begin{split} & \Phi_{\text{LS}} = \arg \min_{\Phi \in \mathbb{D}_{*}^{N}, \|\Phi^{-1}\|_{F}=1} \|\Phi^{-1}Y - SX_{LS}(\Phi)\|_{F} \\ & = \arg \min_{\Phi \in \mathbb{D}_{*}^{N}, \|\Phi^{-1}\|_{F}=1} \|\Phi^{-1}Y - SS^{\dagger}\Phi^{-1}Y\|_{F} \\ & = \arg \min_{\Phi \in \mathbb{D}_{*}^{N}, \|\Phi^{-1}\|_{F}=1} \|Q\Phi^{-1}Y\|_{F} \\ & = \arg \min_{\Phi \in \mathbb{D}_{*}^{N}, \|\Phi^{-1}\|_{F}=1} \|\text{vec}(Q\Phi^{-1}Y)\|_{2} \end{split}$$
(39)

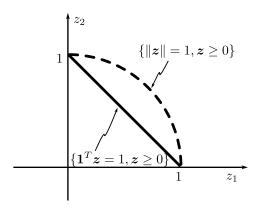


Fig. 3. Illustrating the convexification of the feasible set of interest in \mathbb{R}^2 . The original nonconvex feasible set, depicted by the dashed arc, is replaced with the chord connecting the two endpoints of the arc.

$$= \arg \min_{\mathbf{\Phi} = \operatorname{diag}(z \odot^{-1}), \|z\|_{2} = 1} \left\| \begin{bmatrix} \mathbf{Q} \mathbf{D}_{1} \\ \vdots \\ \mathbf{Q} \mathbf{D}_{J} \end{bmatrix} \mathbf{z} \right\|_{2}$$

$$= \arg \min_{\mathbf{\Phi} = \operatorname{diag}(z \odot^{-1}), \|z\|_{2} = 1} \|A\mathbf{z}\|_{2}$$

$$= \arg \min_{\mathbf{\Phi} = \operatorname{diag}(z \odot^{-1}), \|z\|_{2} = 1} \|A^{T} A \mathbf{z}\|_{2}$$

$$= \operatorname{diag} \left(\mathbf{z}_{\min} \odot^{-1} \right), \tag{40}$$

where the operator $\text{vec}(\cdot)$ in (39) stacks columns of a matrix into a big column vector; (40) follows from Lemma 3 and the definition of z_{\min} . Substituting (40) into (38) yields

$$X_{\rm LS} = S^{\dagger} \operatorname{diag}(z_{\rm min}) Y$$
,

completing the proof.

With the aid of Thm. 3, Alg. 1 finds not only the exact solution in the noiseless case but also the least squares solution of problem (P1). We want to recall, however, that the diagonal entries of Φ are the albedos of the object, and so must be all positive. It follows that the vector $z = \varphi \odot^{-1}$ must be pointwise positive as well. Therefore the positivity constraint on z should also be incorporated into the recovery. In order to do so, we first note that, finding the eigenvector corresponding to the smallest eigenvalue of M is nothing but solving the optimization problem

$$\min \|Mz\|_2 \quad \text{s.t.} \quad \|z\|_2 = 1. \tag{41}$$

Now we can adjust (41) by adding the positivity constraint as

$$\min \|Mz\|_2 \quad \text{s.t.} \quad \|z\|_2 = 1 \text{ and } z \ge 0. \tag{42}$$

Solving the optimization problem (42) is hard due to the nonconvexity of the feasible set (See Fig. 3). The easiest way to convexify this problem is to replace the arc $\{\|z\|_2 = 1, z \ge 0\}$ on the unit sphere with its chord $\{\mathbf{1}^T z = 1, z \ge 0\}$. This means the problem is modified as

$$\min \|Mz\|_2$$
 s.t. $\mathbf{1}^T z = 1$ and $z \ge 0$. (43)

Solving (43) is now easy using some well-developed convex programming method. The above modification of Alg. 1 is summarized in Alg. 2 under the name optimization-based factorization.

Algorithm 2 Optimization-Based Factorization

Inputs: Y, S

Outputs: Estimates $\hat{\Phi}$, \hat{X} of Φ , X

- 1. Compute $M = (I SS^{\dagger}) \odot (YY^{T})$.
- 2. Use a convex programming method to find z_* that minimizes $\|Mz\|_2$ s.t. $\mathbf{1}^T z = 1$, $z \ge 0$.
- 3. Return $\hat{\boldsymbol{\Phi}} = \operatorname{diag}(\boldsymbol{z}_* \odot^{-1}); \ \hat{\boldsymbol{X}} = \boldsymbol{S}^{\dagger} \operatorname{diag}(\boldsymbol{z}_*) \boldsymbol{Y}.$

B. No Approximation Error

If there is no approximation error, i.e. $W_1 = \mathbf{0}$, (35) simplifies to

$$Y = \Phi SX + W_2.$$

and so (P0) becomes

$$\min_{\mathbf{\Phi} \in \mathbb{D}_{+}^{N}, X \in \mathbb{R}^{M \times J}} \|Y - \mathbf{\Phi} SX\|_{F}.$$
 (P2)

Following the same strategy for solving (P1), we first fix Φ and solve for X as

$$X_{LS}(\mathbf{\Phi}) = (\mathbf{\Phi}S)^{\dagger}Y = ((\mathbf{\Phi}S)^{T}(\mathbf{\Phi}S))^{-1}(\mathbf{\Phi}S)^{T}Y$$
$$= (S^{T}\mathbf{\Phi}^{2}S)^{-1}S^{T}\mathbf{\Phi}Y. \tag{44}$$

Substituting into (P2) we get

$$\mathbf{\Phi}_{\mathrm{LS}} = \min_{\mathbf{\Phi} \in \mathbb{D}_{*}^{N}} \| \mathbf{Y} - \mathbf{\Phi} \mathbf{S} (\mathbf{S}^{T} \mathbf{\Phi}^{2} \mathbf{S})^{-1} \mathbf{S}^{T} \mathbf{\Phi} \mathbf{Y} \|_{F}.$$
 (45)

This nonlinear least squares problem seems to be intractable. Instead of solving (45), we can heuristically solve for Φ and X using alternating least squares. In particular, (44) solves for X in terms of Φ . Conversely, if X is fixed, then $\Phi_{LS}(X) = \operatorname{diag}(\hat{\varphi})$ can be solved element-wise as

$$\hat{\varphi}_i = \arg\min_{\varphi_i \in \mathbb{R}} \|\mathbf{y}_i^T - \varphi_i \mathbf{s}_i^T \mathbf{X}\|_2 = \frac{\mathbf{s}_i^T \mathbf{X} \mathbf{y}_i}{\mathbf{s}_i^T \mathbf{X} \mathbf{X}^T \mathbf{s}_i}, \tag{46}$$

for i = 1, ..., N, where s_i^T and y_i^T are i-th rows of S and Y respectively. If we insist in the positivity of albedos, it is clear that (46) should be modified as

$$\hat{\varphi}_i = \arg\min_{\varphi_i \ge \varepsilon} \|\mathbf{y}_i^T - \varphi_i \mathbf{s}_i^T \mathbf{X}\|_2 = \left[\frac{\mathbf{s}_i^T \mathbf{X} \mathbf{y}_i}{\mathbf{s}_i^T \mathbf{X} \mathbf{X}^T \mathbf{s}_i} \right]_{\perp}, \quad (47)$$

where ε is an arbitrarily small positive number, and $[\cdot]_+$ is a short notation for $\max(\cdot, \varepsilon)$. Putting (44) and (47) together, we arrive at the so called alternating factorization, described in Alg. 3.

VII. COMPUTATIONAL ISSUES

Implementation of Algs. 1 and 2 is costly due to the large size of the SVD/optimization problem. The size N of vector z is equal to the number of pixels in an image which is typically large. Nonetheless, we can reduce the size of the problem significantly by observing that

$$Y_{\mathcal{J}} = \Phi_{\mathcal{J}} S_{\mathcal{J}} X, \quad \forall \mathcal{J} \subset \mathcal{N}, \tag{48}$$

Algorithm 3 Alternating Factorization

Inputs: Y, S

Outputs: Estimates $\hat{\Phi}$, \hat{X} of Φ , X.

1. Initialize $\hat{\Phi} = I_N$.

2. Set
$$\hat{\boldsymbol{X}} = (\boldsymbol{S}^T \hat{\boldsymbol{\Phi}}^2 \boldsymbol{S})^{-1} \boldsymbol{S}^T \hat{\boldsymbol{\Phi}} \boldsymbol{Y}$$
.

3. Set
$$\hat{\varphi}_i = \left[\frac{\mathbf{s}_i^T \hat{\mathbf{X}} \mathbf{y}_i}{\mathbf{s}_i^T \hat{\mathbf{X}} \hat{\mathbf{X}}^T \mathbf{s}_i}\right]_+, \quad i = 1, \dots, N.$$

4. Repeat steps 2 and 3 until improvement in mean squared error is negligible.

where $\Phi_{\mathcal{J}} = \operatorname{diag}(\{\varphi_i\}_{i\in\mathcal{J}})$. Now $\Phi_{\mathcal{J}}$ and X can be found by solving a similar SVD/optimization problem of size $|\mathcal{J}| \ll N$. Once X is found, the full matrix Φ can be recovered elementwise as in Alg. 3. If the equation is exact, the subset \mathcal{J} can be chosen arbitrarily as long as $S_{\mathcal{J}}$ has NSFR. Recall from Proposition 3 that $S_{\mathcal{J}}$ being NSFR also implies that S is NSFR. However, if it is noisy, a careful selection of \mathcal{J} for stable recovery has to be done. The following result characterizes the deviation of a singular vector associated with the smallest singular value of a matrix from the actual solution of the corresponding homogeneous equation under some perturbation. This will serve as a guidance for choosing a "good" subset \mathcal{J} .

Proposition 4: Suppose A is a matrix with condition number $\kappa(A)$ and δA is a perturbation such that $A + \delta A$ has a nontrivial null space. Let z and $z + \delta z$ be unit norm singular vectors associated with the smallest singular values of A and $A + \delta A$, respectively. Then we have the following bound

$$\|\delta z\|_2 \lesssim 1 + 2\kappa^2(A) \frac{\|\delta A\|_2}{\|A\|_2},$$
 (49)

where the approximate inequality \lesssim assumes that $(\delta A)^T \delta A \approx 0$.

Proof: Let $\sigma_1(A)$ and $\sigma_n(A)$ be the largest and smallest singular values of A respectively. z being the singular vector associated with the smallest singular value of A means

$$\mathbf{A}^T \mathbf{A} \mathbf{z} = \sigma_n^2(\mathbf{A}) \mathbf{z}. \tag{50}$$

Since $A + \delta A$ has nontrivial null space, its smallest singular value must be 0, and so

$$(A + \delta A)(z + \delta z) = \mathbf{0},$$

which is also equivalent to

$$(\mathbf{A} + \delta \mathbf{A})^{T} (\mathbf{A} + \delta \mathbf{A})(z + \delta z) = \mathbf{0}.$$
 (51)

Ignoring the second order term $(\delta A)^T \delta A$, we can rearrange terms of (51) and take the norm of both sides to get

$$||A^{T}A\delta z||_{2} = ||A^{T}Az + (A^{T}\delta A + (\delta A)^{T}A)(z + \delta z)||_{2}$$

$$= ||\sigma_{n}^{2}(A)z + (A^{T}\delta A + (\delta A)^{T}A)(z + \delta z)||_{2}$$

$$\leq \sigma_{n}^{2}(A)||z||_{2} + ||(A^{T}\delta A + (\delta A)^{T}A)(z + \delta z)||_{2}$$

$$\leq \sigma_{n}^{2}(A) + ||A^{T}\delta A + (\delta A)^{T}A||_{2} \cdot ||z + \delta z||_{2}$$

Algorithm 4 RANSAC-Based Factorization

Inputs: Y, S, \bar{N}, L

Outputs: Estimates $\hat{\Phi}$, \hat{X} of Φ , X

- 1. Randomly generate L subsets of length \bar{N} of N.
- 2. For each subset \mathcal{J} , compute

$$oldsymbol{M}_{(\mathcal{J})} = (oldsymbol{I} - oldsymbol{S}_{\mathcal{J}} oldsymbol{S}_{\mathcal{J}}^{\dagger}) \odot (oldsymbol{Y}_{\mathcal{J}} oldsymbol{Y}_{\mathcal{J}}^T).$$

- 3. Choose \mathcal{J} with largest $\rho(\mathcal{J})$ as defined in (54).
- 4. Use Alg. 2 to solve for $\hat{\Phi}_{\mathcal{J}}$, \hat{X} from $Y_{\mathcal{J}}$, $S_{\mathcal{J}}$.

5. Set
$$\hat{\varphi}_i = \left[\frac{s_i^T \hat{X} y_i}{s_i^T \hat{X} \hat{X}^T s_i}\right]_+, \quad i = 1, \dots, N.$$

6. Return \hat{X} in step 4 and $\hat{\Phi} = \operatorname{diag}(\hat{\varphi}_1, \dots, \hat{\varphi}_N)$.

$$\leq \sigma_n^2(A) + \|A^T \delta A\|_2 + \|(\delta A)^T A\|_2$$

$$\leq \sigma_n^2(A) + 2\|A\|_2 \|\delta A\|_2$$

$$= \sigma_n^2(A) + 2\sigma_1^2(A) \frac{\|\delta A\|_2}{\|A\|_2}.$$

On the other hand

$$\|\mathbf{A}^T \mathbf{A} \delta \mathbf{z}\|_2 \ge \sigma_n^2(\mathbf{A}) \|\delta \mathbf{z}\|_2. \tag{52}$$

Hence

$$\|\delta z\|_{2} \le 1 + 2\frac{\sigma_{1}^{2}(A)}{\sigma_{n}^{2}(A)} \cdot \frac{\|\delta A\|_{2}}{\|A\|_{2}} = 1 + 2\kappa^{2}(A)\frac{\|\delta A\|_{2}}{\|A\|_{2}}.$$
 (53)

Back to the issue of choosing a "good" subset \mathcal{J} , for a particular \mathcal{J} , we want to solve for the eigenvector associated with the smallest eigenvalue of the matrix $M_{(\mathcal{J})} = (I - I)$ $S_{\mathcal{J}}S_{\mathcal{J}}^{\dagger})\odot(Y_{\mathcal{J}}Y_{\mathcal{J}}^{T})$. On the one hand, Proposition 4 suggests that in order to have a stable solution, $\mathcal J$ should be chosen so that the condition number of $M_{(\mathcal{J})}$ is small. On the other hand, we also want the nullity of $M_{(\mathcal{J})}$ to be numerically close to 1 in some sense so that the eigenvector associated with the smallest eigenvalue is close to the real solution. One heuristic way to measure this closeness is to compute the drop from the second smallest eigenvalue $\lambda_{N-1}(M_{\mathcal{J}})$ to the smallest eigenvalue $\lambda_N(M_{\mathcal{J}})$ relatively to the largest eigenvalue $\lambda_1(M_{\mathcal{I}})$ of $M_{(\mathcal{I})}$. Thus, we heuristically combined these two observations to define a single quantity associated with a subset \mathcal{J} to measure the numerical rank-(N-1) of $M_{(\mathcal{J})}$ as

$$\rho(\mathcal{J}) = \frac{1}{\kappa(\mathbf{M}_{\mathcal{J}})} \cdot \frac{\lambda_{N-1}(\mathbf{M}_{\mathcal{J}}) - \lambda_{N}(\mathbf{M}_{\mathcal{J}})}{\lambda_{1}(\mathbf{M}_{\mathcal{J}})}$$
$$= \frac{\lambda_{N}(\mathbf{M}_{\mathcal{J}})\lambda_{N-1}(\mathbf{M}_{\mathcal{J}}) - \lambda_{N}^{2}(\mathbf{M}_{\mathcal{J}})}{\lambda_{1}^{2}(\mathbf{M}_{\mathcal{J}})}. \tag{54}$$

We will search for \mathcal{J} with largest $\rho(\mathcal{J})$ over all subsets of \mathcal{N} . However, this leads to an NP combinatorial optimization problem. We propose instead an algorithm to choose the best \mathcal{J} in terms of $\rho(\mathcal{J})$ over just a relatively small number of random subsets of \mathcal{N} . Alg. 4 is similar, to some extent, to the RANSAC (RANdom SAmple Consensus) [8], a commonly used algorithm in computer vision for model fitting. We will therefore call it RANSAC-based factorization.

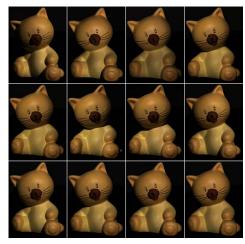


Fig. 4. Dataset 1: 12 images of a real object under different directional light

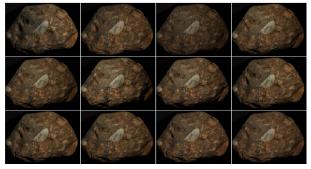


Fig. 5. Dataset 2: 12 images of a real object under different directional light sources.

VIII. EXPERIMENTAL RESULTS

A. Dealing With Color Images

For inverse rendering from color images, we need to solve (13) for the three color channels simultaneously to eliminate the scaling ambiguity of the albedos across channels. Specifically, let Y_R , Y_G , Y_B be the image matrices for red, green, and blue channels respectively. Also, let φ_R , φ_G , φ_B be the albedo vectors associated with red, green, and blue channels respectively. Then the actual matrix factorization we want implement is given by the concatenating equation

$$\begin{bmatrix} Y_{R} \\ Y_{G} \\ Y_{B} \end{bmatrix} = \operatorname{diag} \begin{pmatrix} \begin{bmatrix} \varphi_{R} \\ \varphi_{G} \\ \varphi_{B} \end{bmatrix} \end{pmatrix} \cdot \begin{bmatrix} S \\ S \\ S \end{bmatrix} X. \tag{55}$$

Forming this equation is important for implementations, but the equation itself is mathematically the same as (13) and does not change any of the above theoretical results.

B. Simulations

Simulations are performed on two real datasets,⁴ each of which has J=12 images of an object as shown in Figs. 4 and 5, using MATLAB with the convex programming cvx provided by [28]. The images were taken under different directional light sources. The light directions of the

⁴The data and code for photometric stereo are available at: http://pages.cs.wisc.edu/~csverma/CS766 09/Stereo/stereo.html.

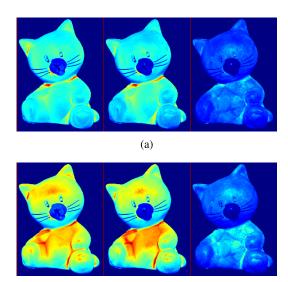


Fig. 6. Dataset 1: Reconstruction of albedo maps in red, green and blue channels from left to right using either (a) alternating or (b) RANSAC-based factorization.

(b)

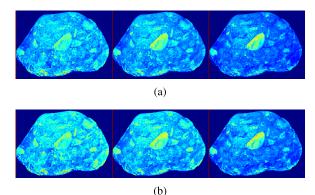


Fig. 7. Dataset 2: Reconstruction of albedo maps in red, green and blue channels from left to right using either (a) alternating or (b) RANSAC-based factorization.

12 sources can be calibrated and will be used as a ground truth for the inverse rendering.

The normal map of the object was estimated from the images and light directions using photometric stereo. We want to note that the light directions were only used to estimate the geometry. In our inverse rendering, the lighting, however, is assumed to be unknown and arbitrary. Of course, in practice we are aiming at estimating the geometry directly using some depth-based 3D modeling framework. The spherical harmonic matrix S was then computed from the surface normals of the object using (9) and (11), with M = 9. The size of each image is 340×512 ; excluding zero-intensity pixels results in N = 35983 for dataset 1, and N = 72561 for dataset 2. We can easily confirm that the spherical harmonic matrix Sindeed has NSFR for both datasets by using Proposition 3 and Corollary 1. In particular, we randomly chose an index subset $\mathcal J$ and check if the rank of the matrix $(I-S_{\mathcal J}S_{\mathcal J}^{\scriptscriptstyle \intercal})\odot$ $(S_{\mathcal{J}}S_{\mathcal{J}}^T)$ is equal to $|\mathcal{J}|-1$.

The albedo and lighting matrices $\hat{\Phi}$ and \hat{X} were estimated using either Alg. 3 or 4. The reconstructed albedo maps for 3 color channels using both methods are visualized in Figs. 6 and 7 using the jet color map. To quantify the

TABLE I

DATASET 1: ANGLES BETWEEN THE RECONSTRUCTIONS OF LIGHT SOURCE DIRECTIONS AND THEIR GROUND TRUTHS USING DIFFERENT METHODS: ALTERNATING AND RANSAC-BASED FACTORIZATIONS. THE MEAN ERRORS OVER ALL 12 LIGHT SOURCES ARE 6.8 DEGREES FOR ALTERNATING AND 2.7 DEGREES FOR RANSAC-BASED

source	1	2	3	4	5	6	7	8	9	10	11	12
Alternating	9.6	4.9	1.2	8.6	8.1	8.9	7.9	5.6	7.5	7.5	3.0	8.5
RANSAC	6.3	6.3	1.6	3.6	1.7	3.4	2.1	1.6	0.4	1.0	2.0	2.1

TABLE II

DATASET 2: ANGLES BETWEEN THE RECONSTRUCTIONS OF LIGHT SOURCE DIRECTIONS AND THEIR GROUND TRUTHS USING DIFFERENT METHODS: ALTERNATING AND RANSAC-BASED FACTORIZATIONS. THE MEAN ERRORS OVER ALL 12 LIGHT SOURCES ARE 9.9 DEGREES FOR ALTERNATING AND 1.8 DEGREES FOR RANSAC-BASED

source	1	2	3	4	5	6	7	8	9	10	11	12
Alternating	18.0	7.4	12.0	9.5	7.7	9.8	10.5	11.2	9.1	8.7	8.4	7.0
RANSAC	4.5	0.9	4.3	0.8	1.7	0.8	0.8	1.5	1.1	0.2	1.6	3.1

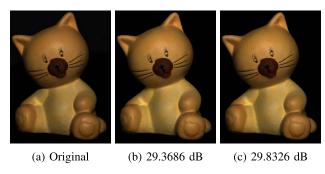


Fig. 8. Dataset 1: Reconstruction of one of the original images (a) using either alternating (b) or RANSAC-based factorization (c) with corresponding SNRs. The average SNRs over all given twelve images are 24.9786 dB for alternating and 24.8493 dB for RANSAC-based.



Fig. 9. Dataset 2: Reconstruction of one of the original images (a) using either alternating (b) or RANSAC-based factorization (c) with corresponding SNRs. The average SNRs over all given twelve images are 27.1829 dB for alternating and 27.4583 dB for RANSAC.

accuracy of our model, the twelve original images in Y were compared in SNR (Signal-to-Noise Ratio) to the reconstructed images by forming $\hat{Y} = \hat{\Phi} S \hat{X}$. The results are shown in Figs. 8 and 9 for one of the images in each dataset.

Although both methods can reconstruct the original images reasonably well, the reconstruction error itself does not quantify the accuracy of estimating albedos and lighting individually. Therefore we did another step to estimate the light directions for which we have ground truths. In particular, from (9), the x, y, z components of the light directions can be extracted from rows 2, 3, 4 of the lighting matrix X, respectively. The angle between each recovered light direction and its ground truth was then computed for all 12 sources and shown in Tabs. I and II. It can be seen that the RANSAC-based

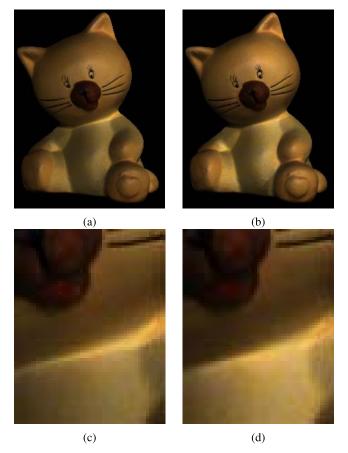


Fig. 10. Dataset 1: Random relighting using (a) alternating and (b) RANSAC-based factorization. The magnifications of an interconnection region are shown respectively below in (c) and (d).

is much more accurate than the alternating factorization, with a mean error of 2.7 degrees versus 6.8 degrees for dataset 1, and 1.8 degrees versus 9.9 degrees for dataset 2.

For the purpose of relighting, new images were rendered by randomly generating a novel lighting matrix X_{new} and computing $Y_{\text{new}} = \hat{\Phi} S X_{\text{new}}$. The relit images using alternating and RANSAC-based factorizations are demonstrated in Figs. 10 and 11. It can be subjectively observed from Fig. 10 that relighting using the alternating



Fig. 11. Dataset 2: Random relighting using (a) alternating and (b) RANSAC-based factorization.

factorization behaves poorly at the interconnection regions of the object where our convex assumption is violated. It is because the measurement noise model that the algorithm is based on does not account for approximation error resulting from the nonconvexity of the object. The difference between the two relighting methods is less visible in Fig. 11 because the 'rock' object does not have abrupt transitions like the neck of the 'cat' object. We remark that the implementations of the alternating algorithm were faster. In contrast, relighting using RANSAC-based factorization, at the cost of increasing the computations, produces very realistic looking images that are well-behaved at the interconnection regions.

IX. CONCLUDING REMARKS

We have studied the inverse rendering problem of a Lambertian convex object with distant light sources. Under these assumptions, all possible images of the object live close to a low-dimensional linear subspace spanned by the first few spherical harmonics. The inverse rendering thus becomes a factorization of the image matrix into a product of a diagonal albedo and a lighting matrix with a known SH matrix in the middle. This special matrix factorization is unique if and only if the SH matrix associated with the object has nonseparable full rank, a stronger notion of full rank. In the noiseless case, the exact solution (up to some scale) can be found via an SVD-based algorithm. When an approximation error is present, the SVD-based algorithm also yields the least squares solution to the factorization. An optimization-based algorithm is introduced to enforce the positivity of albedos. Because the number of pixels in each images is typically large, a heuristic RANSAC-based algorithm is proposed to reduce the size of the optimization problem. When a measurement noise is present in the model, an alternating algorithm is introduced instead for the factorization.

APPENDIX A PROOF OF PROPOSITION 2

A. KFR⇒NSFR

Suppose \mathcal{N}_1 and \mathcal{N}_2 are nonempty disjoint subsets such that $\mathcal{N}_1 \cup \mathcal{N}_2 = \mathcal{N}$. We will show that $\operatorname{rank}(S_{\mathcal{N}_1}) + \operatorname{rank}(S_{\mathcal{N}_2}) \geq M + 1$. Indeed, since S has full Kruskal rank, any k rows of S are linearly independent for all $k \leq M$. Thus, $\operatorname{rank}(S_{\mathcal{N}_1}) = \min\{M, |\mathcal{N}_1|\}$, and $\operatorname{rank}(S_{\mathcal{N}_2}) = \min\{M, |\mathcal{N}_2|\}$.

It follows that

where (a) follows from $|\mathcal{N}_1| + |\mathcal{N}_2| = |\mathcal{N}| = N$.

$B. NSFR \Rightarrow SFR$

For $i \in \mathcal{N}$, since **S** has NSFR, it must be that

$$rank(S_{\{i\}}) + rank(S_{\mathcal{N}\setminus\{i\}}) \ge M + 1.$$

However, since S contains no zero rows, it is clear that $\operatorname{rank}(S_{\{i\}}) = 1$. Hence, $\operatorname{rank}(S_{\mathcal{N}\setminus\{i\}}) \geq M$. On the other hand, $\operatorname{rank}(S_{\mathcal{N}\setminus\{i\}}) \leq \operatorname{rank}(S) = M$. Therefore, $\operatorname{rank}(S_{\mathcal{N}\setminus\{i\}}) = M$, or S has strong full rank.

C. $SFR \Rightarrow FR$

If S has strong full rank, then for i=1 we have $\operatorname{rank}(S) \geq \operatorname{rank}(S_{N\setminus\{1\}}) = M$. Since $S \in \mathbb{R}^{N\times M}$, we also have $\operatorname{rank}(S) \leq M$. Thus, $\operatorname{rank}(S) = M$, or S has full rank.

Now consider the reverse implications. If N = M + 1, it is obvious that the definitions of strong full rank and Kruskal full rank coincide. Thus, SFR \Rightarrow KFR \Rightarrow NSFR. Finally, for $N \ge M + 2$ we can show that the reverse implications are not true by constructing simple counterexamples. We leave that to the readers due to the space constraint. For a visual proof, see Figure 2.

APPENDIX B PROOF OF PROPOSITION 3

Let \mathcal{N}_1 and \mathcal{N}_2 be nonempty disjoint subsets of \mathcal{N} such that $\mathcal{N} = \mathcal{N}_1 \cup \mathcal{N}_2$. We need to show that $\operatorname{rank}(S_{\mathcal{N}_1}) + \operatorname{rank}(S_{\mathcal{N}_2}) \ge M + 1$. Indeed, consider the following cases:

If $\mathcal{J} \subset \mathcal{N}_1$ then

$$rank(S_{\mathcal{N}_1}) + rank(S_{\mathcal{N}_2}) \ge rank(S_{\mathcal{J}}) + rank(S_{\mathcal{N}_2})$$
$$= M + rank(S_{\mathcal{N}_2}) \ge M + 1.$$

If $\mathcal{J} \subset \mathcal{N}_2$, similarly we also have $\operatorname{rank}(S_{\mathcal{N}_1}) + \operatorname{rank}(S_{\mathcal{N}_2}) \geq M + 1$.

If \mathcal{J} is not a subset of either \mathcal{N}_1 or \mathcal{N}_2 , put $\mathcal{J}_1 = \mathcal{J} \cap \mathcal{N}_1$, and $\mathcal{J}_1 = \mathcal{J} \cap \mathcal{N}_1$. It is evident that $\mathcal{J}_1, \mathcal{J}_2$ are nonempty disjoint subsets of \mathcal{J} such that $\mathcal{J} = \mathcal{J}_1 \cup \mathcal{J}_2$. Since $S_{\mathcal{J}}$ has nonseparable full rank, it must be that

$$\operatorname{rank}(\boldsymbol{S}_{\mathcal{N}_1}) + \operatorname{rank}(\boldsymbol{S}_{\mathcal{N}_2}) \geq \operatorname{rank}(\boldsymbol{S}_{\mathcal{J}_1}) + \operatorname{rank}(\boldsymbol{S}_{\mathcal{J}_2}) \geq M + 1.$$

Hence, S has NSFR, completing the proof.

APPENDIX C PROOF OF LEMMA 1

For all $\mathcal{I} \subset \mathcal{N}$, we can write

$$\hat{S}_{\mathcal{I}} = [S_{\mathcal{I}}|(\Psi S)_{\mathcal{I}}] \stackrel{(a)}{=} [S_{\mathcal{I}}|S_{\mathcal{I}}\tilde{X}X^{\dagger}],$$

where (a) follows from (18). Note that every column of $S_{\mathcal{I}}\tilde{X}X^{\dagger}$ is a linear combination of columns of $S_{\mathcal{I}}$. This implies that $\mathcal{R}(\hat{S}_{\mathcal{I}}) = \mathcal{R}(S_{\mathcal{I}})$. Hence

$$\operatorname{rank}(\hat{\mathbf{S}}_{\mathcal{I}}) = \operatorname{rank}(\mathbf{S}_{\mathcal{I}}), \quad \forall \mathcal{I} \subset \mathcal{N}.$$

APPENDIX D PROOF OF LEMMA 2

We first show that $\{[s_i^T | \psi_i s_i^T]^T\}_{i \in \mathcal{B}}$ are linearly independent. Suppose $\{a_i\}_{i \in \mathcal{B}}$ satisfies

$$\sum_{i \in \mathcal{B}} \alpha_i [\mathbf{s}_i^T | \psi_i \mathbf{s}_i^T]^T = \mathbf{0},$$

which implies $\sum_{i \in \mathcal{B}} \alpha_i s_i = \sum_{i \in \mathcal{B}} \alpha_i \psi_i s_i = \mathbf{0}$. By the independence of $\{s_i\}_{i \in \mathcal{B}}$, it must be that $\alpha_i = 0, \forall i \in \mathcal{B}$, or $\{[s_i^T | \psi_i s_i^T]^T\}_{i \in \mathcal{B}}$ are linearly independent.

Next, it follows from Lemma 1 that

$$\dim\left(\mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{I}}^T)\right) = \dim\left(\mathcal{R}(\boldsymbol{S}_{\mathcal{I}}^T)\right) = |\mathcal{B}|.$$

It means that $\{[s_i^T | \psi_i^T s_i^T]^T\}_{i \in \mathcal{B}}$ is a set of linearly independent vectors in the space $\mathcal{R}(\hat{S}_{\mathcal{I}}^T)$ whose cardinality is equal to the dimension of the space. Thus, it is a basis for $\mathcal{R}(\hat{S}_{\mathcal{I}}^T)$.

APPENDIX E

PROOF OF THE NECESSITY PART OF THEOREM 1

Suppose that S does not have NSFR. We will construct a solution $(\tilde{\Phi}, \tilde{S})$ to (13) such that either $\tilde{\Phi}$ is not a scaled version of Φ_0 or \tilde{X} is not a scaled version of X_0 . Consider the following cases.

Case 1: S has a zero row, say the first row. It follows that the first row of Y is also zero. We can therefore simply choose $\tilde{X} = X_0$ and $\tilde{\Phi} = \text{diag}(2, 1, ..., 1) \odot \Phi_0$.

Case 2: S does not have full column rank. It means that columns of S are linearly dependent. There thus exists a matrix \tilde{X} that is not a scaled version of X_0 such that $S\tilde{X} = SX_0$. We then choose (Φ_0, \tilde{X}) as a solution.

Case 3: S has full column rank with no zero rows. It follows from the definition of NSFR that there exists a nontrivial subset \mathcal{J} of \mathcal{N} such that $\operatorname{rank}(S_{\mathcal{J}}) + \operatorname{rank}(S_{\mathcal{J}^c}) = M$. Let $\Psi = \operatorname{diag}(\psi)$, where $\psi_j = 2, \forall j \in \mathcal{J}$ and $\psi_j = 1, \forall j \in \mathcal{J}^c$. Consider the matrix $\hat{S} = [S|\Psi S]$.

We first show that $\operatorname{rank}(\hat{\boldsymbol{S}}_{\mathcal{J}}) = \operatorname{rank}(\boldsymbol{S}_{\mathcal{J}})$. Let $\{\boldsymbol{s}_k\}_{k\in\mathcal{K}}$ be a basis for $\mathcal{R}(\boldsymbol{S}_{\mathcal{J}}^T)$, we can show that $\{[\boldsymbol{s}_k^T|2\boldsymbol{s}_k^T]^T\}_{k\in\mathcal{K}}$ is a basis for $\mathcal{R}(\hat{\boldsymbol{S}}_{\mathcal{J}}^T)$. Indeed, similarly to the proof of Lemma 2, those vectors are independent. Moreover, any row of $\hat{\boldsymbol{S}}_{\mathcal{J}}$ can be written as

$$[s_j^T | 2s_j^T] = \sum_{k \in \mathcal{K}} \alpha_k [s_k^T | 2s_k^T], \quad \forall j \in \mathcal{J},$$
 (56)

where $s_j = \sum_{k \in \mathcal{K}} \alpha_k s_k$ is the linear representation of s_j in terms of $\{s_k\}_{k \in \mathcal{K}}$. This implies that $\{[s_k^T | 2s_k^T]^T\}_{k \in \mathcal{K}}$ is a basis for $\mathcal{R}(\hat{S}_{\mathcal{J}}^T)$, and so $\operatorname{rank}(\hat{S}_{\mathcal{J}}) = \operatorname{rank}(S_{\mathcal{J}}) = |\mathcal{K}|$. Similarly, $\operatorname{rank}(\hat{S}_{\mathcal{J}^c}) = \operatorname{rank}(S_{\mathcal{J}^c})$. Therefore

$$\operatorname{rank}(\hat{S}) \leq \operatorname{rank}(\hat{S}_{\mathcal{J}}) + \operatorname{rank}(\hat{S}_{\mathcal{J}^c})$$

$$= \operatorname{rank}(S_{\mathcal{J}}) + \operatorname{rank}(S_{\mathcal{J}^c}) = M. \tag{57}$$

On the other hand, if $\{s_i\}_{i\in\mathcal{B}}$ is a basis for $\mathcal{R}(S^T)$, then similarly to the proof of Lemma 2, $\{[s_i|\psi_is_i]\}_{i\in\mathcal{B}}$ are linearly independent. It means that the dimension of $\mathcal{R}(\hat{S}^T)$ is at least $|\mathcal{B}|$, or

$$rank(\hat{S}) \ge rank(S) = M. \tag{58}$$

Combining (57) and (58) yields $\operatorname{rank}(\hat{S}) = \operatorname{rank}(S) = M$. It follows that each column of the matrix ΨS must be a linear combination of columns of S. In other words, there exists a matrix Γ such that

$$\Psi S = S\Gamma. \tag{59}$$

Choose $\tilde{\Phi} = \operatorname{diag}(\tilde{\varphi})$, where $\tilde{\varphi_j} = \varphi_j/2, \forall j \in \mathcal{J}$ and $\tilde{\varphi_j} = \varphi_j, \forall j \in \mathcal{J}^c$. Left-multiplying by $\tilde{\Phi}$ and right-multiplying by X_0 to both sides of (59) we get

$$\mathbf{\Phi}_0 \mathbf{S} \mathbf{X}_0 = \tilde{\mathbf{\Phi}} \mathbf{S} \tilde{\mathbf{X}},\tag{60}$$

where $\tilde{X} = \Gamma X_0$. It shows that $(\tilde{\Phi}, \tilde{X})$ is a solution to (13). However, by the construction, $\tilde{\Phi}$ is not a scaled version of Φ .

APPENDIX F PROOF OF LEMMA 3

Let $Q = I - SS^{\dagger}$. It is easy to check that Q is a projection matrix, i.e., $Q^2 = Q = Q^T$. Since D_i is a diagonal matrix for all j, we have

$$(\boldsymbol{Q}\boldsymbol{D}_j)^T = \boldsymbol{D}_j^T \boldsymbol{Q}^T = \boldsymbol{D}_j \boldsymbol{Q}, \quad j = 1, \dots, J.$$

Hence

$$\boldsymbol{A}^T = [\boldsymbol{D}_1 \boldsymbol{Q} | \boldsymbol{D}_2 \boldsymbol{Q} | \cdots | \boldsymbol{D}_K \boldsymbol{Q}].$$

It follows that

$$A^{T}A = \sum_{i=1}^{J} D_{j} Q^{2} D_{j} = \sum_{i=1}^{J} D_{j} Q D_{j}.$$
 (61)

Thus for any $i, k \in \{1, 2, ..., N\}$, we get

$$(\mathbf{A}^{T}\mathbf{A})_{ik} = \sum_{j=1}^{J} \sum_{m,n=1}^{N} (\mathbf{D}_{j})_{im} (\mathbf{Q})_{mn} (\mathbf{D}_{j})_{nk}$$

$$= \sum_{j=1}^{J} (\mathbf{D}_{j})_{ii} (\mathbf{Q})_{ik} (\mathbf{D}_{j})_{kk}$$

$$= (\mathbf{Q})_{ik} \sum_{j=1}^{J} (\mathbf{Y})_{ij} (\mathbf{Y})_{kj}$$
(62)

$$= (\mathbf{Q})_{ik} \sum_{j=1}^{J} (\mathbf{Y})_{ij} (\mathbf{Y}^{T})_{jk}$$

 $= (\boldsymbol{Q})_{ik} (\boldsymbol{Y} \boldsymbol{Y}^T)_{ik},$

where (62) follows from the diagonality of D_i s and (63) follows from the definition of D_i s. This implies that $A^T A = Q \odot (YY^T)$. Substituting Q by $I - SS^{\dagger}$ completes the proof.

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